

TERMODINAMIKAREN ZERO PRINTZIPIOA

TENPERATURA

1.5

* OREKA TERMIKOA

- (SISTEMA, NORMA, INGURUNEA) TIURKOTEAREN KASU BEREZIA
ADIABATIKOA ISOLAMENDUA
DIATERMANOA

1.5

* TERMODINAMIKAREN ZERO PRINTZIPIOA

1.6

* TENPERATURA KONZEPtua

- ESPERIENTZIAN OINARRITURIKO METODOA : PROPIETATE KOMUNA
- MATEMATIKAN OINARRITURIKO METODOA : BALIO KOMUNA
 - LERRO ISOTERMANOAK
 - LERRO ISOTERMANO KORRESPONDENTEAK
 - TENPERATURA :

1.7

* SISTEMEN TENPERATURA

1.8

- TENPERATURA ESKALA

1.9

- TERMOMETROA

1.10

- SISTEMA-MOTA

1.11

- SISTEMA KONKRETUA

- ALDAGAI TERMOMETRIKOA

- FUNTzio TERMOMETRIKOA

- PROPORZIONALTASUN-KONSTANTEA : PUNTU FIJKOAK

- TENPERATURAREN NEURKETA

- TERMOMETRO-MOTAK

- TENPERATURA-ESKALAK

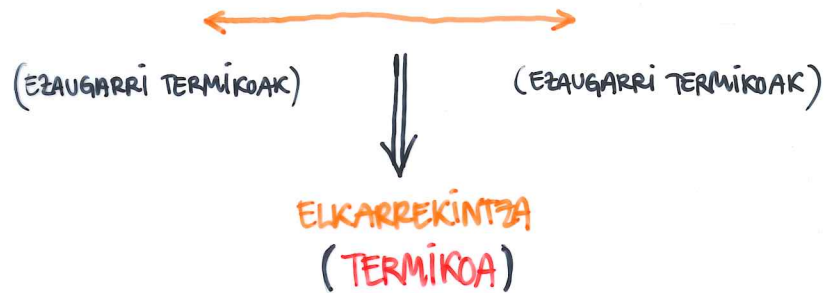
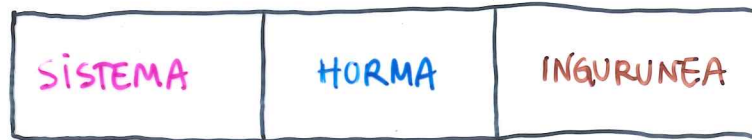
- GAS IDEALAREN TENPERATURA-ESKALA

- CELSIUS ESKALA

- KELVIN ESKALA

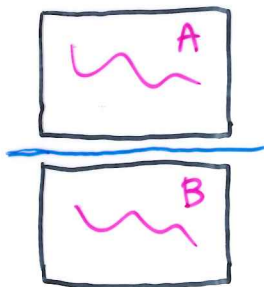
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OREKA TERMÍKOA



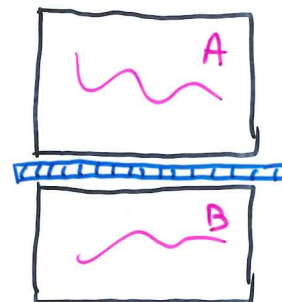
BAI

HORMA DIATERMIKOA
ISOLAMENDU EZA TERMÍKOA



EZ

HORMA ADIABATÍKOA
ISOLAMENDUA TERMÍKOA



OREKA TERMÍKOA :

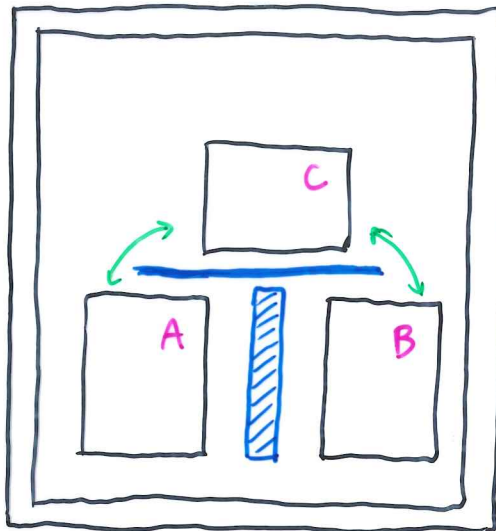
HORMA DIATERMIKOAAREN BIDEZ 'LOTURÍKO' BI SISTEMEK LORTUKO DUTEN
EGOERA ALDAEZINA (HOTTS; OREKA-EGOERA)

- Aldagai termodinamikak
aipatu; $\{X, Y, Z, \dots\}$
ez dute biren balioak
aldatuko; edo, bai.

TERMODINAMIKAREN ZERO PRINTZIPIOA

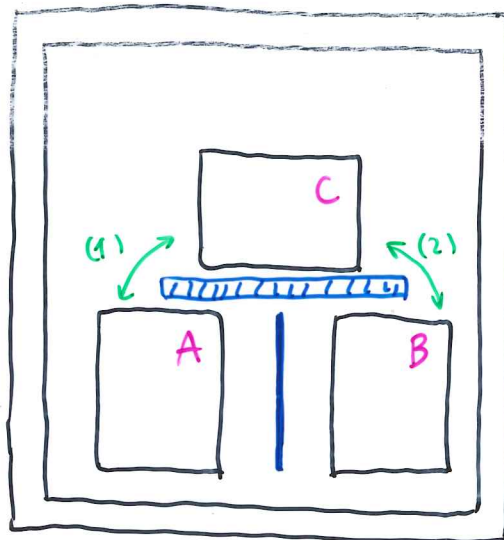
(7,8 orr.)

(a)



BIAK ALDIBEREAN

(b)

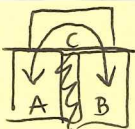


LEHENENGO BAT,
GERO BESTEA

TERMODINAMIKAREN ZERO PRINTZIPIOA :

ESKU ARTEAN DITUGUN A ETA B SISTEMAK, C SISTEMAREKIN OREKA TERMIKUAN BADAUDE (BAKUTZA BEHE ALDETIK), ELKARREN ARTERO OREKA TERMIKUAN DAUDE

(i)



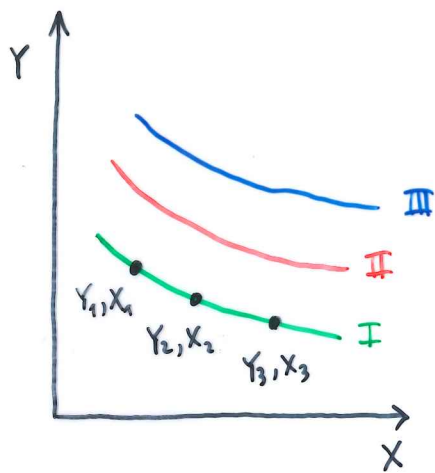
propietate
transitiboa

(ii)

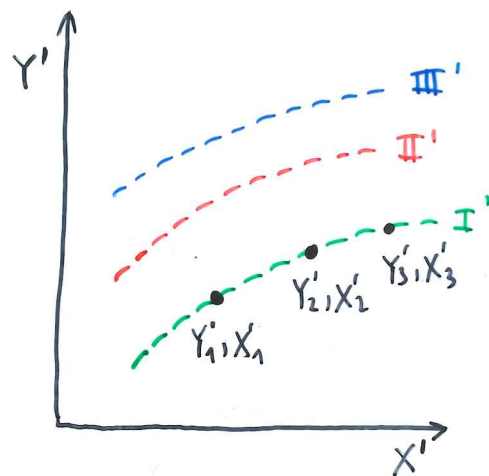


TENPERATURA

(a) ESPERIENTZIAN OINARRITUTAKO "METODOA" **TENPERATURA** DEFINITZEKO



A SISTEMA



B SISTEMA

"**LERRO**" ISOTERMANOA

LERRO ISOTERMANO KORRESPONDENTEAK

TENPERATURA: PROPIETATE KOMUNA

(b) MATEMATIKAN OINARRITUTAKO "METODOA" **TENPERATURA** DEFINITZEKO
 $\{A, B, C\}$ SISTEMAK

$\{A, C\}$ OREKA TERMIKOA

$$f_{AC}(X_A, Y_A; X_C, Y_C) = 0$$

$\{B, C\}$ OREKA TERMIKOA

$$f_{BC}(X_B, Y_B; X_C, Y_C) = 0$$

$$f_{AC}(X_A, Y_A; X_C, Y_C) = 0 \Rightarrow Y_C = g_{AC}(X_A, Y_A; X_C)$$

$$f_{BC}(X_B, Y_B; X_C, Y_C) = 0 \Rightarrow Y_C = g_{BC}(X_B, Y_B; X_C)$$

$$g_{AC}(X_A, Y_A; X_C) = g_{BC}(X_B, Y_B; X_C)$$

$\{A, B\}$ OREKA TERMIKOA

$$f_{AB}(X_A, Y_A; X_B, Y_B) = 0$$

$$g_{ABC}(X_A, Y_A; X_B, Y_B; X_C) = 0$$

X_C SOBERAN

$$h_A(X_A, Y_A) = h_B(X_B, Y_B)$$

$$h_A(X_A, Y_A) = h_B(X_B, Y_B) = h_C(X_C, Y_C)$$

$$\text{TENPERATURA : } t = h_A(X_A, Y_A) = h_B(X_B, Y_B) = h_C(X_C, Y_C)$$

OROKORTUZ ...

$$h_i = h_i(x_i, y_i, z_i, \dots)$$

$$i = A, B, C \dots$$

$$t = h_A(x_A, y_A, z_A, \dots) = \dots = h_i(x_i, y_i, z_i, \dots) = \dots$$

SISTEMAK

ALDAGAIK

FUNTZIOAK

} DESBERDINAK

(SISTEMA-MOTA ERE DESBERDINA IZAN DAITE)

OREKA TERMIKOAN BADAUDE

"TEMPERATURA BEREAN DAUDE"

FUNTZIOAK BALIO BAKARRA DU (BALIO KOMUNA DU)

HAUXE DA LERRO ISOTERMANOAREN SORTARI BALIOA

ZEIN BALIO DA HORI? (ZEIN ZENBAKI DA HORI?)

BALIOAK FINKATEKO TEMPERATURA-ESKALA DEFINITU BEHAR DA

TEMPERATURA-ESKALA / TEMPERATURAREN NEURKETA

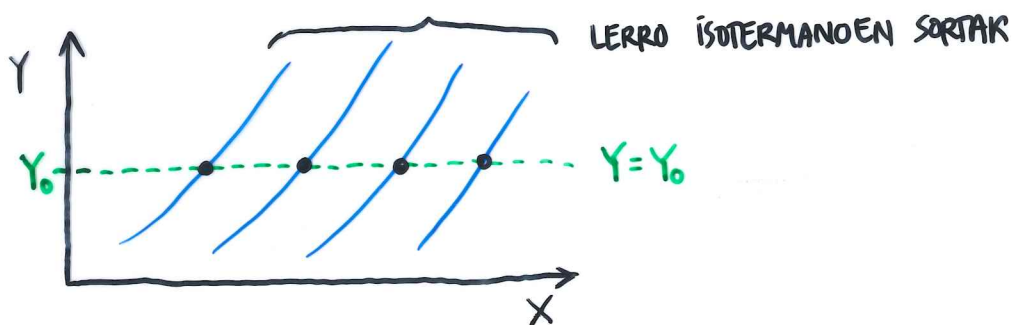
- BALIOA FINKATREKO PROZEDURA OROKORRA

- | | | |
|---------------------------|-----------|-------------------------|
| 1 - SISTEMA-MOTA | } FINKATU | } TERMOMETROA (PATROIA) |
| 2 - SISTEMA KONKRETUA | | |
| 3 - ALDAGAI TERMOMETRIKOA | | |
| 4 - FUNTzio TERMOMETRIKOA | | |
| 5 - PUNTU FINKOAK | | |

1 - S SISTEMA-MOTA $S = S(X, Y, Z, \dots)$ $\{X, Y, Z, \dots\}$

2 - SISTEMA KONKRETUA $S(X, Y)$

3 -



X DA ALDAGAI TERMOMETRIKOA $\{X; Y_0\}$
(HAUKE DA TEMPERATURAREKIN BATERA ALDATUKO DENA)

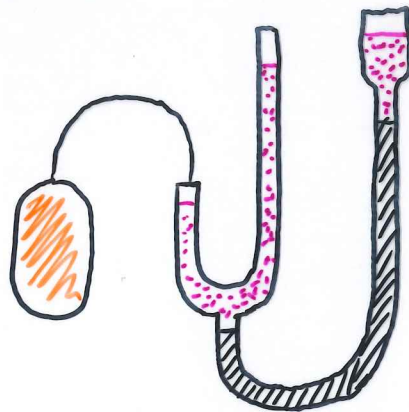
4 - $\theta(X) = aX$ FUNTzio TERMOMETRIKOA $a?$

5 - PUNTU FINKOAK ARBITRARIOKI FINKATU X_0 FINKATU $\theta_0 = \theta(X_0)$ } a FINKATU $a = \frac{\theta_0}{X_0}$

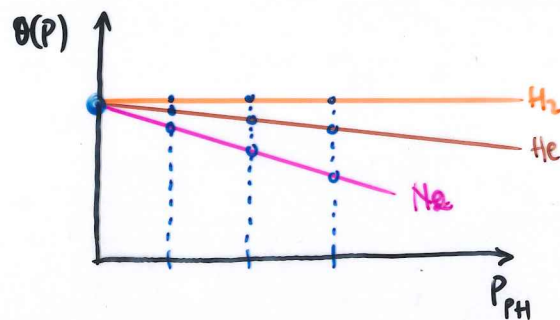
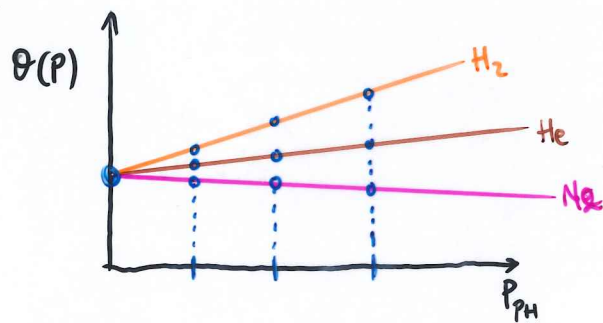
$$\theta = \left[\frac{\theta_0}{X_0} \right] X$$

GAS IDEALAREN TEMPERATURA-ESKALA

(a) TRESNA

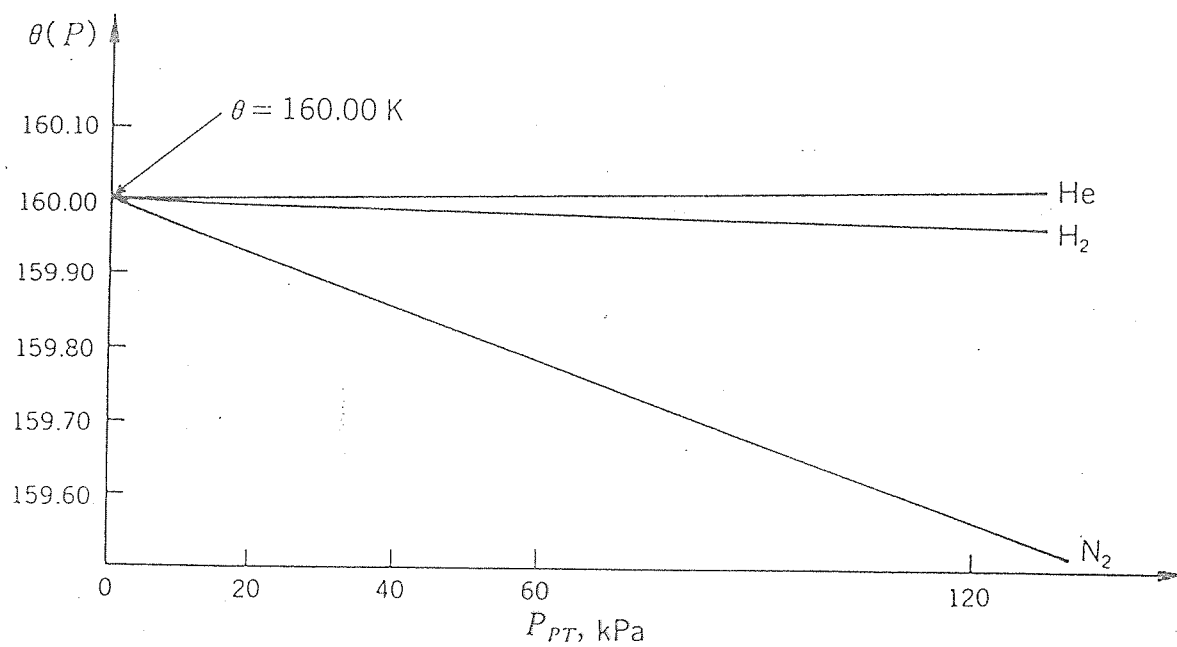
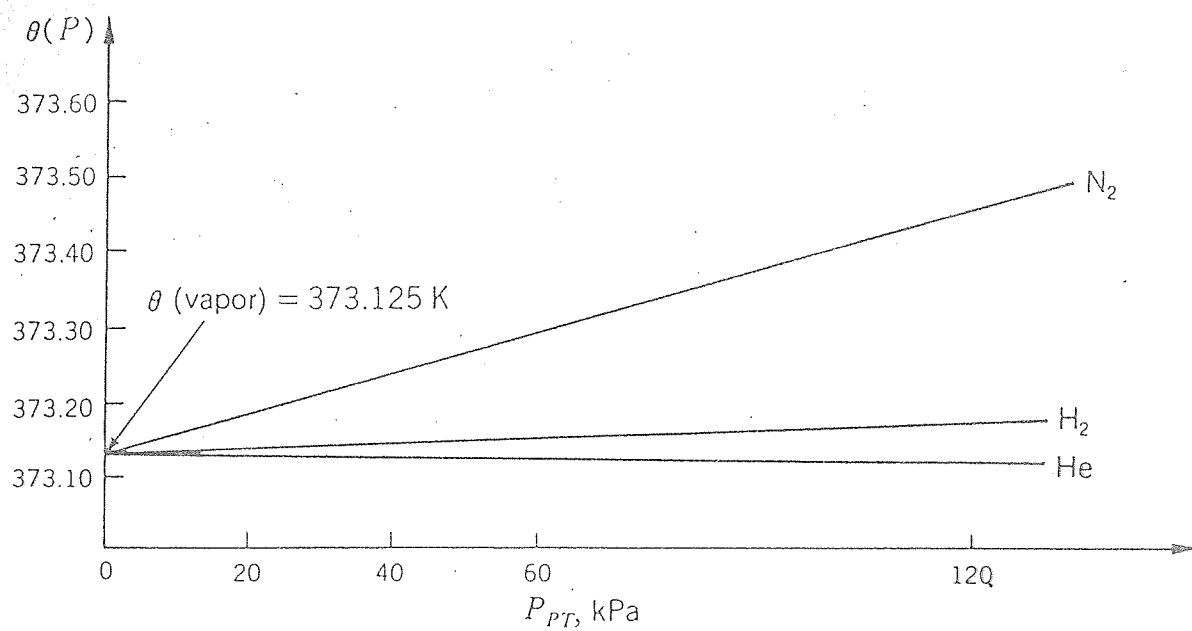


(b) PROZEDURA



$$\theta = 273.16 \lim_{P_{PH} \rightarrow 0} \left(\frac{P}{P_{PH}} \right) \quad V = K_{te}$$

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En segundo lugar, β es un parámetro más natural para expresar la temperatura que la propia T . Según veremos más tarde, el cero absoluto de temperatura ($T = 0$) no puede alcanzarse en un número finito de pasos, lo que puede resultar confuso; es menos sorprendente que un valor infinito de β (el que tiene β cuando $T = 0$) no sea alcanzable en un número finito de pasos. Sin embargo, aunque β sea la manera más natural de expresar temperaturas, tal terminología es incómoda en su uso diario. No resulta fácil decir que el agua, que se congela a 0°C (273 K), lo hace cuando $\beta = 2,65 \times 10^{20}\text{ J}^{-1}$, o que el punto de ebullición, a 100°C (373 K), se alcanza cuando $\beta = 2,47 \times 10^{20}\text{ J}^{-1}$. Como tampoco lo son los valores que caracterizan un día frío (10°C , esto es, $\beta = 2,56 \times 10^{20}\text{ J}^{-1}$) u otro más cálido (20°C , es decir, $\beta = 2,47 \times 10^{20}\text{ J}^{-1}$).

En tercer lugar, la existencia y el valor de la constante fundamental k es mera consecuencia de que insistamos en utilizar una escala convencional de temperaturas en vez de la más fundamental basada en β . Las escalas Fahrenheit, Celsius y Kelvin no van bien encaminadas: el inverso de la temperatura, esencialmente β , es una medida de la temperatura con mayor significado, más natural. No debe esperarse, sin embargo, que sea aceptada como tal, por motivos históricos y por el poder que en nuestra cultura tienen los números simples, como 0 y 100, e incluso 32 y 212, mucho más apropiados, además, para el día a día.

Aunque la constante de Boltzmann k suele incluirse en la lista de constantes fundamentales, esto no es más que una consecuencia de una mala elección histórica. Si Ludwig Boltzmann hubiera llevado a cabo sus trabajos antes que Fahrenheit y Celsius los suyos, se habría visto que β era la medida natural para la temperatura, y podríamos habernos acostumbrado a expresar la temperatura en unidades del inverso del julio, con valores bajos de β para sistemas calientes y altos para sistemas fríos.

Sin embargo, la convención estaba ya establecida con los sistemas calientes a temperaturas más altas que los fríos, y se introdujo k , mediante la relación $k\beta = 1/T$, para hacer corresponder la escala natural de temperaturas basada en β con la escala basada en T , convencional y profundamente arraigada. Esto es, la constante de Boltzmann no es más que un factor de conversión entre una escala convencional y bien establecida y la escala que, visto con perspectiva, la sociedad podría haber adoptado. La constante de Boltzmann no habría sido necesaria si β hubiera sido adoptada como medida de la temperatura.

Terminaremos esta sección con un detalle positivo. Hemos establecido que la temperatura, y concretamente β , es un parámetro que da cuenta de la distribución en el equilibrio de las moléculas de un sistema entre los estados disponibles de energía. Uno de los sistemas más fáciles de imaginar en este contexto es un gas perfecto (o «ideal»), en el que imaginamos las moléculas como integrantes de un enjambre caótico, algunas de ellas moviéndose rápidamente, otras despacio, viajando en línea recta hasta chocar entre sí, rebotando en direcciones diferentes y a velocidades distintas, y golpeando contra las paredes en una tormenta de impactos, dando lugar así a lo que interpretamos como presión. Un gas es un conjunto caótico de moléculas (de hecho, las palabras «gas» y «caos» provienen de la misma raíz), caótico en la distribución espacial y en la distribución de velocidades moleculares. Cada velocidad se corresponde con una energía cinética dada, y se puede utilizar la distribución de Boltzmann para expresar la distribución de velocidades, a través de la distribución de moléculas entre los estados posibles de energía de traslación; y relacionar esa distribución de velocidades con la temperatura. La expresión resultante se denomina *distribución de velocidades de Maxwell-Boltzmann*, ya que fue James Clerk Maxwell (1831-1879) el primero que la dedujo, de una forma ligeramente diferente. Cuando se lleva a cabo el cálculo,

The Meaning of Temperature

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The summer of 1988 found me working with a dozen high-school teachers on the subject "Atoms in Motion." One of our classes was devoted to "the meaning of temperature." The topic may seem simple, but it can be descriptive. Misconceptions are propagated in many textbooks, and there are several surprises. What follows is an expanded version of the notes I handed out following our class.

The Essence

Among the many things that I learned from Professor Eric Rogers was a simple, clear meaning for temperature: temperature is hotness measured on some definite scale.¹ In turn, "hotness" is the tendency to transfer energy in irregular, microscopic fashion. (It is the tendency to transfer energy "as heat.")

We can order objects according to their hotness. A red-hot piece of iron is hotter than iron we can hold in our hands; it is not as hot as white-hot iron. If we place the three pieces in contact, as sketched in Fig. 1, then energy is transferred from "red hot" to "can be handled" and to "red hot" from "white hot." "Temperature" provides such an ordering on a convenient numerical scale. In our illustration, the temperatures might be these

$$T_{\text{can be handled}} = 300 \text{ K} \quad (1)$$

$$T_{\text{red hot}} = 1000 \text{ K} \quad (2)$$

$$T_{\text{white hot}} = 1700 \text{ K} \quad (3)$$

To say that object A is "hotter" than object B is to say that if A and B are placed together, then energy flows from A to B.

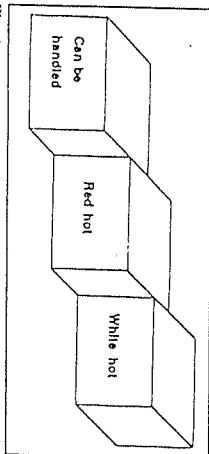


Fig. 1. Among the three objects, energy is transferred from right to left. That sequence provides an ordering of the objects according to their "hotness."

As time goes on, the rates of energy transfer between our blocks decrease toward zero. When the rates become zero, we have thermal equilibrium.² The three blocks now have equal hotnesses; that is, equal tendencies to transfer energy, which cancel out one another's effects. Equal

hotnesses imply equal temperatures. In short, mutual thermal equilibrium implies equality of temperatures and vice versa.

While correct, this presentation of what temperature means is austere and can leave one with a desire for a more intuitive grasp. The attempt to provide a more intuitive view, however, may lead to error, as we will see.

A Common Misconception

For a dilute gas, the empirical gas law gives the pressure p as

$$p = \frac{1}{V} N k T \quad (4)$$

where N is the total number of molecules, V is the container's volume, k is Boltzmann's constant, and T is the absolute temperature.

The kinetic theory of gases provides its own expression for the pressure

$$p = \frac{1}{3} \frac{N}{V} m v^2 \quad (5)$$

where m is the molecular mass and v is the molecular speed (specifically, the root mean square speed).

We can equate the two expressions for the pressure and then rearrange factors to isolate the kinetic energy

$$\frac{1}{2} m v^2 = \frac{3}{2} k T \quad (6)$$

While a true statement, the equation does not provide us with a meaning for temperature. We may not take the equation to imply that "absolute temperature means the average kinetic energy (within a constant factor)."³ There are at least two reasons why the equation does not provide an acceptable meaning for temperature.

① Equation (6) is strictly a classical result. When the gas is no longer dilute or the temperature is low, quantum physics becomes essential to a correct description. (Intermolecular forces become significant, too, but that is a separate issue.) When quantum physics just begins to be important, Eq. (6) must be replaced by

$$\frac{1}{2} m v^2 = \frac{3}{2} k T \left[1 + C \frac{N}{V} \left(\frac{h^2}{2\pi m k T} \right)^{3/2} \right] \quad (7)$$

Here h denotes Planck's constant, and the constant C depends on the species of gas.^{2,3} For example, for ^4He , the usual helium isotope with two neutrons, the constant C is $-(2^{-3/2})$; for ^3He , the rare isotope with only one neutron, the constant is $+(2^{-3/2})$. The average kinetic energy depends on number density N/V as well as on

temperature, and the dependence on T is complicated.

(Editor's Note: For 1 mole of ^4He , compressed to 1 liter at 10 K, the correction factor is 0.2 percent.) We would not be tempted to use this more exact equation to ascribe a meaning to temperature based on average kinetic energy. (Moreover, note this: in a mixture of ^3He and ^4He gases, with each gas specified to have the same number density N/V and certainly having the same temperature, the average kinetic energies would differ between the two isotopes because the constants C differ in sign. At the same temperature, you can have different average kinetic energies—surely temperature cannot mean average kinetic energy.)

② When a kitchen oven is baking brownies, it is filled with radiation, mostly infrared radiation. The energy density of that radiation grows as the fourth power of the absolute temperature

$$\left(\frac{\text{energy per unit volume}}{\text{of radiation}} \right) = \left(\frac{\text{universal}}{\text{constant}} \right) T^4 \quad (8)$$

We can solve this equation for T

$$T = \left(\frac{\text{energy density}}{\text{universal constant}} \right)^{1/4} \quad (9)$$

But we would not be tempted to say, "Oh, now I know what temperature means. It is the fourth root of the radiant energy density."

Just as we do not use Eq. (8) or (9) to give a meaning to temperature, so we should not use Eq. (6) to try to give a meaning. All three equations are valid instances in which temperature plays a role in determining physical properties related to energy, but we should not push the equations beyond that.

Some Thermal Surprises

Negative absolute temperatures. Experiments can produce temperatures below absolute zero: negative absolute temperatures. The first such experiment was performed by Edward Purcell and Robert Pound in 1951; their physical system was the assembly of nuclei in a crystal of lithium fluoride.^{4,5} (In more detail, the system was the nuclei acting as though they were tiny spinning bar magnets. That aspect of the nuclei was adequately decoupled from the nuclei acting as though they were vibrating masses, having ordinary kinetic energy; the latter aspect cannot exhibit a negative temperature.) By swiftly reversing an external magnetic field and achieving internal magnetic thermal equilibrium before a second reversal occurred, Purcell and Pound produced a final temperature of -350 K .

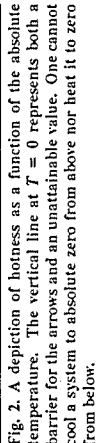
(A technical aside. One would like some microscopic picture of how the negative temperature was achieved. The interactions of the magnetic moments among themselves are crucial. In a crystal, a magnetic moment interacts not only with an external magnetic field but also with the neighboring magnetic moments. Following convention, we call the energy associated with the latter interaction the spin-spin energy. Initially, the spin system is at a

positive temperature, with the nuclear magnetic moments predominantly parallel to the external field, an orientation of negative potential energy. The first field reversal is so fast that the spins cannot respond; they cannot change their orientation during the reversal. Consequently, the spins find themselves pointed predominantly antiparallel to the reversed field, an orientation of positive potential energy. In a short time some of this energy is transferred, by mutual interactions, to spin-spin energy. The energy associated with the mutual interactions becomes relatively large. In fact, that energy becomes larger than what one would predict at any positive temperature.

Since the interaction between the spin system and the lattice vibrations is weak, the spin system holds onto that relatively large amount of energy, even as the crystal is transferred for observation from the small external field to a large magnetic field. There is, in fact, so much spin-spin energy around that the magnetic moments "prefer" to line up antiparallel to the new external field, a position of positive potential energy, rather than line up in the usual parallel position, the position of negative potential energy. There is enough energy around so that a spin can easily line up in what is ordinarily an energetically unfavored orientation. This occurs no matter how the crystal, as a macroscopic object, is rotated during the physical transfer from one magnetic field in the laboratory to the other.

Indeed, the examination of the crystal in the large field was quite directly a test for the amount of energy possessed by the nuclear spin system. Purcell and Pound applied electromagnetic radiation at a frequency chosen so that lithium-7 nuclei could flip from an orientation of low to high potential energy by the absorption of a photon. Instead of absorption, they found—as they had hoped—the stimulated emission of radiation at that frequency. The spin system had so much energy that spins predominantly flipped to low potential energy orientations and simultaneously emitted a photon. From the direct observation of stimulated emission of radiation—rather than absorption—one infers both that the total magnetic moment is antiparallel to the external field and that the nuclear spin system is at a negative temperature.)

Negative T is hotter than positive T . At $T = -350 \text{ K}$, Purcell and Pound's system of nuclei was hotter than anything at a positive temperature. That is, it would transfer energy to anything at a positive temperature. The experimental test for a negative temperature relied on that property: the nuclei now emitted energy when probed with radio-frequency waves. At ordinary room temperature, they had absorbed such waves. Indeed, all negative temperatures are hotter than all positive temperatures.^{6,7} You cannot get to zero from either side. As their system cooled down, its temperature went from -350 K to -1000 K to $-\infty$, which is physically equivalent to $+\infty$, and continued through $+1000 \text{ K}$ to $+350 \text{ K}$, ordinary room temperature. The system did not pass through absolute zero.



Bizarre results, indeed, and there is yet another surprise.

is. A bicycle pump provides an example wherein the air is compressed by the plunger. Specify that no heating (as with a Bunsen burner or electric current) accompanies the compression. (Thus, in technical language, we are examining adiabatic compression.) The mechanical work associated with the compression increases the physical system's energy. As a consequence of all of this, does the temperature increase? Most often the temperature does increase, but – amazingly – sometimes it decreases.

Another technical aside. The nuclei in the solid phase of ^3He are located at specific lattice sites, and so in some respects the nuclei act independently of each other. Their spins need not be correlated by the Pauli exclusion principle. Nuclei in the liquid phase do not enjoy this quasi-independence, and so the Pauli principle (or antisymmetry or a total wave function) causes the spins to be correlated and reduces their contribution to the entropy.

The solid phase has a smaller volume per atom than does the liquid phase. Thus, when a coexisting mixture of liquid and solid is compressed adiabatically, solid is

Although this description does not provide all the microscopic explanation that one would like to see, it does outline the logic of why the reversible, adiabatic compression of coexisting phases of ^4He causes the temperature to decrease.]

Recapitulation

Many texts, including some of the very best, contain statements like this: "Temperature" is a measure of the average kinetic energy of the molecules or atoms in a substance."¹⁰ Such a statement is misleading, if not downright wrong. The function of temperature is not to tell us about a physical system's *amount* of energy. Rather, it is to tell us about a system's *tendency to transfer energy* (as *temperature*). To be sure, a quantitative knowledge of the *temperature* to transfer, together with some statistical reasoning, sometimes enables us to infer the amount of energy. Equations (6), (7), and (8) are cases in point. We reason from *temperature* to energy. None of those equations, however, captures the essential meaning of temperature. For that we must return to "temperature is hotness measured on some definite scale."

1. E. M. Rogers, *Physics for the Inquiring Mind* (Princeton University Press, Princeton, NJ, 1960), p. 412.
2. Ralph Baierlein, *Atoms and Information Theory* (Freeman, San Francisco, 1971), p. 307.
3. L. D. Landau and E. M. Lifshitz, *Statistical Physics* (Addison-Wesley, Reading, MA, 1958), pp. 158–159.
4. E. M. Purcell and R. V. Pound, *Phys. Rev.* **81**, 279 (1951).
5. J. H. Van Vleck, *Nuovo Cimento Supplement* **6** (10), 1081 (1957).
6. Norman F. Ramsey, *Phys. Rev.* **103**, 20 (1956).
7. Mark W. Zemansky, *Temperatures Very Low and Very High* (Van Nostrand, Princeton, NJ, 1964), p. 112.
8. D. S. Betts, *Refrigeration and Thermometry Below One Kelvin* (Susssex University Press, England, 1976), p. 52.
9. K. Mendelssohn, *The Quest for Absolute Zero*, 2nd ed. (Halden-Wiley, New York, 1977), p. 267.
10. Paul G. Hewitt, *Conceptual Physics, Teacher's Edition* (Addison-Wesley, Reading, MA, 1987), p. 301.

[illegible]

Podría parecer como si el principio de conservación de la energía dictara que disponemos de toda la energía que queremos. La naturaleza no gana ni pierde energía. Sin embargo, la energía del mar, por poner un ejemplo, el movimiento térmico de todos los átomos del mar, está prácticamente fuera de nuestro alcance. Con el fin de organizar y dominar esta energía, de poder utilizarla, tenemos que poseer una

El principio de conservación de la energía indica que la energía total del mundo se mantiene constante. Pero debido a los movimientos irrregulares de las partículas, esta energía puede distribuirse de manera no uniforme que, en determinadas circunstancias, no sea posible extraer más energía de la que se añade, es decir, que no haya forma de controlarla.

Creo que mediante una analogía podré dar cuenta mejor de esta dificultad. No sé si alguno de ustedes habrá pasado por la experiencia, que yo conozco bien, de estar sentado en la playa con varias personas, cuando de pronto empieza a caer una tempestad. Rápidamente los garramos las llavias e corremos a cubierto. Una vez protegidos de la lluvia empezamos a reanarnos y descubrimos que a la vuelta está un poco lloviendo, pero relativamente no hasta que nuestro cuerpo, con lo que vamos desahogando con ella hasta que llega un momento en que está demasiado empapada, es decir, que muy pronto desahogamos nos moja. Probamos con la siguiente talla y muy pronto descubrimos algo terrible, que todas las llavias están tan mojadas como nuestro cuerpo. Ya no hay pues, manera de secarse un poco más a diferencia de poder varias toallas, porque en algún sentido ya no hay diferencia entre cada una de las toallas que estamos usando y las toallas que están mojadas. Podría inventar una expresión de cantidad y llamarla "índice de secar el agua". La toalla posee la misma facilidad de secar el agua que la toalla que no, con la cual cada vez que paso la toalla por encima de mí, me moje, tanto agua pasa de la toalla a mí como de mí a la toalla.

Esto no significa que haya la misma cantidad de agua en la toalla que sobre mi cuerpo — una toalla grande contendrá más agua que una pequeña — pero tienen la misma humedad. Cuando todo adquiere mismo grado de humedad va no hay nada que hacer.

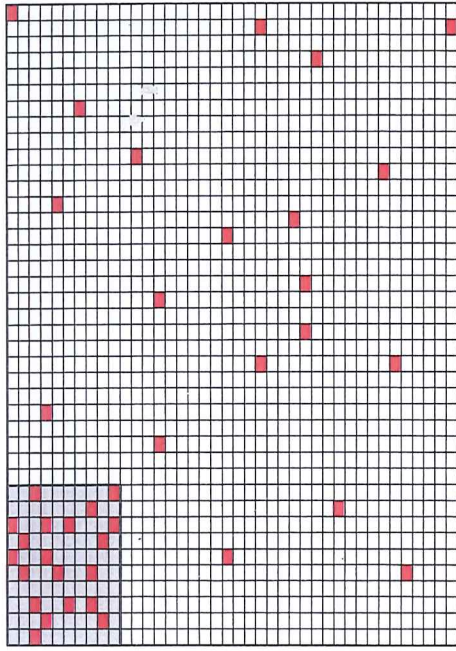
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Así pues, en el caso de la rueda de trinquete con aspas, que es un

OREKA TERMIKOAREN
ADIBIDEA

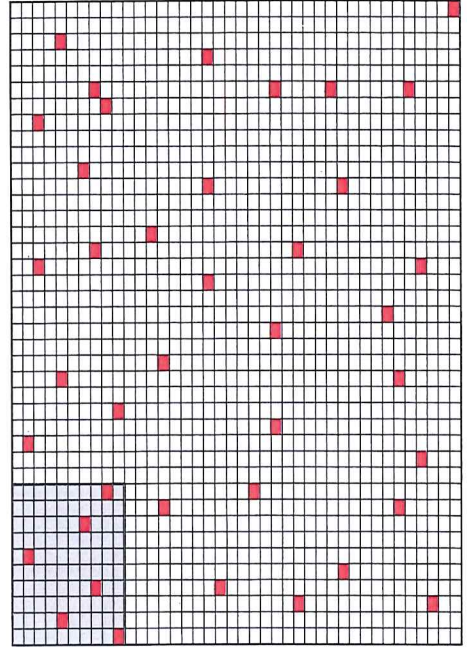
"El carácter de la Ley física", Richard Feynman
CONJETURAS, Antoni Bosch Editor, 1983

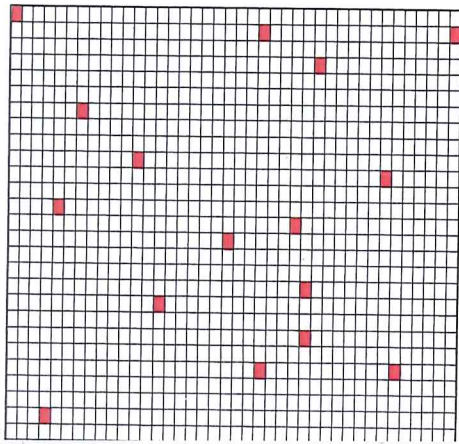
Some time after the initial stage, the energy is spread more uniformly over all the atoms as a result of their jostling each other. The small block still has a higher proportion of its atoms ON than the bigger block, and so it is still hotter. The temperature of System 1 is now 0.72 and that of System 2 is 0.23.



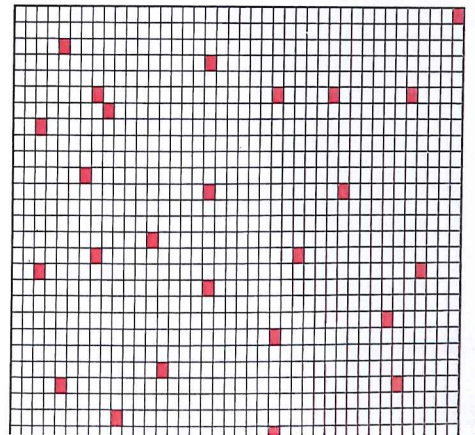
What is the final state of the universe? There is *no* final state for the careful observer, for the ON-ness jostles and migrates forever (there is no rule that brings it to an end). But there is an *apparent* final state for an observer who stands so far back that the behavior of the individual atoms cannot be discerned. There is a final state for the *thermodynamic observer*, not for the atomic individualist. This *apparent* end of change occurs when there is a *uniform distribution of ON-ness*, as in the figure below.

Later, the jostling of the atoms results in a uniform distribution of the energy. There will be small accumulations here and there (there are fluctuations), but on average the proportion of atoms ON in the smaller block is equal to the proportion ON in the larger. The temperatures of the blocks are now the same, at 0.27, and they are at thermal equilibrium.





verse? There is *no* final state for the es and migrates forever (there is no ere is an *apparent* final state for an he behavior of the individual atoms te for the *thermodynamic observer*, not nt end of change occurs when there in the figure below.



The sequence of illustrations on the two preceding pages shows how the universe attains not so much a final state as a *steady state*. In this state the individual atoms turn ON and OFF as they have always done, but, to the casual observer of averages, the redistribution of energy leaves the universe apparently unchanged. We see that the jostling, random migration of energy disperses it. When it is uniformly dispersed over the available universe, it remains dispersed.

That last remark is not quite true, because the random wandering of ON-ness may lead it to accumulate, by chance, in System 1 and leave Sytem 2 completely OFF. However, even with a universe of 1,600 atoms this chance is slight (as may be tested by using one of the programs), and in a real Universe, where each system is a block of Avogadro's numbers of atoms, the chance is so remote that it is negligible. *Lack of rules allied with vastness of domain accounts for the virtual irreversibility of the process of dispersal.*

Temperature

Before we wrap this observation into a neat package, let us notice that we are also closing in on the significance of *temperature*. We have just seen that System 1 heats System 2 as a natural consequence of the dispersal of energy, and that the net transfer of energy continues until, on average, the energy is evenly dispersed over all the available atoms. Now note the following important distinction. When the ON-ness is evenly distributed, there is more *energy* in System 2 than in System 1 (because the former contains more atoms, and therefore more are ON when the ON-ness is uniformly distributed), but the *ratio* of the numbers ON and OFF is the same in both.

All this conforms with common sense about hot and cold so long as we interpret the ratio of the numbers of atoms ON and OFF as indicating *temperature*. First, we know that energy flows as heat from high temperatures to low, and we have seen that System 1 (which initially has a higher "temperature" than System 2) heats System 2. Second, the steady state, when there is no net flow of energy between the two systems, corresponds to their having equal "temperatures," not equal total energies. Finally, "temperature" measures the *incoherent* motion, not the coherent motion, of particles; it is intrinsically a thermodynamic (as distinct from a dynamic) property of systems of many particles. It would be absurd to refer to the temperature of a single particle. When we say that a baseball is warm, we are referring to the excitation of its component particles, not to the whole baseball regarded as a single particle.